

AGR 2022 - Workshop Lectures Schedule May 16-17

<https://atlanticgr2022.ca/>

Day 1: Monday May 16
all times in NL time zone (GMT-2:30)

| Time | Speaker | Title |
|-----------|-----------------------------|---|
| 0945-1000 | Ivan Booth and Hari Kunduri | Welcoming remarks and organizational details |
| 1000-1130 | Carla Cederbaum | Where is the center of mass of an isolated relativistic object? (I) |
| 1230-1330 | Daniel Pook-Kolb | Introduction to black hole apparent horizons in numerical relativity (I) |
| 1345-1445 | Daniel Pook-Kolb | Introduction to black hole apparent horizons in numerical relativity (II) |
| 1530-1630 | William East | How to extract energy from a black hole (I) |
| 1645-1745 | William East | How to extract energy from a black hole (II) |

Day 2: Tuesday May 17
all times in NL time zone (GMT-2:30)

| Time | Speaker | Title |
|-----------|-----------------------------|--|
| 1000-1130 | Carla Cederbaum | Where is the center of mass of an isolated relativistic object? (II) |
| 1230-1330 | Daniel Pook-Kolb | Introduction to black hole apparent horizons in numerical relativity (III) |
| 1400-1500 | William East | How to extract energy from a black hole (III) |
| 1500-1700 | Gather meeting event | link to be updated |

Please note that Professor Cederbaum will be giving two 1.5-hour lectures whereas Professor East and Dr Polk-Kolb will be giving three 1-hour lectures.

Workshop Lectures

Where is the center of mass of an isolated relativistic object?

Carla Cederbaum

Universität Tübingen

In this mini-course, we will study relativistic initial data sets and their total "charges". We will begin by defining initial data sets for relativistic spacetimes, including some examples. Then, we will discuss the asymptotic flatness conditions that ensure that the relativistic system modeled by an initial data set is isolated. The central part of the course will be to review how to define asymptotic/total charges such as energy, mass, linear and angular momentum, and, most prominently, the center of mass of asymptotically flat initial data sets. Our review will include traditional approaches by Arnowitt, Deser, and Misner (ADM) as well as geometric approaches by Huisken and Yau and by myself and Sakovich. We will end by describing open problems in this field of research.

Introduction to black hole apparent horizons in numerical relativity

Daniel Pook-Kolb

Max-Planck-Institut für Gravitationsphysik (Albert Einstein Institute) Hanover

This mini-course will cover the concept of apparent horizons and the techniques we use today to find them in simulations of black hole mergers. We will start with trapped surfaces, trapped regions and marginally outer trapped surfaces (MOTSs) introduced by Penrose and Hawking. The world tubes the MOTSs trace out in time then lead us to the definition of the quasilocal (isolated and dynamical) horizons introduced by Ashtekar, Krishnan and others. We will next discuss numerical methods used to locate MOTSs in initial data and in simulated spacetimes, both with and without additional assumptions about symmetries. Finally, we will have a closer look at recent results, which paint an intricate picture of the merger of two black holes. The stability properties of MOTSs will help in the interpretation of these results, and so we will give some room to the "MOTS stability operator" introduced by Andersson, Mars and Simon.

How to extract energy from a black hole

William East

Perimeter Institute of Theoretical Physics

Although an object falling through a black hole horizon can never escape, remarkably, it is still possible to extract energy from a spinning black hole. We will cover the connection between this fact and black hole thermodynamics, and various ways in which black hole energy extraction can be realized, including the particle process originally proposed by Penrose, and the wave analogue called superradiance. In the final part, we will discuss the superradiant instability, and how this can turn astrophysical black holes into particle detectors for ultralight bosons.